

IASI-NG INSTRUMENT PERFORMANCES STATUS

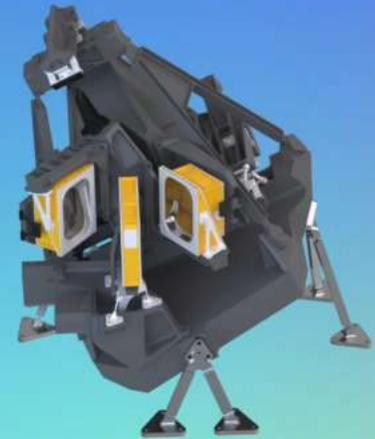
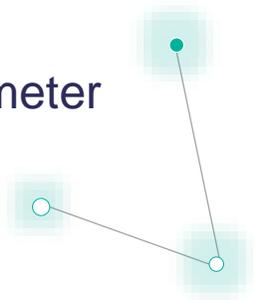
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IASI CONFERENCE 2024

December 02 > 06 2024
NANCY

IASI-NG INSTRUMENT CONCEPT

- Significant performance improvement with wider field of view (factor 4), allowed by innovative instrumental concept proposed by Airbus Defence and Space
- Associated field effects are compensated in pupil by inserting dynamically variable glass thickness
- Achieved through 2 pairs of KBr prisms, synchronized via an unique device movement
- First implementation in space of a Mertz Interferometer



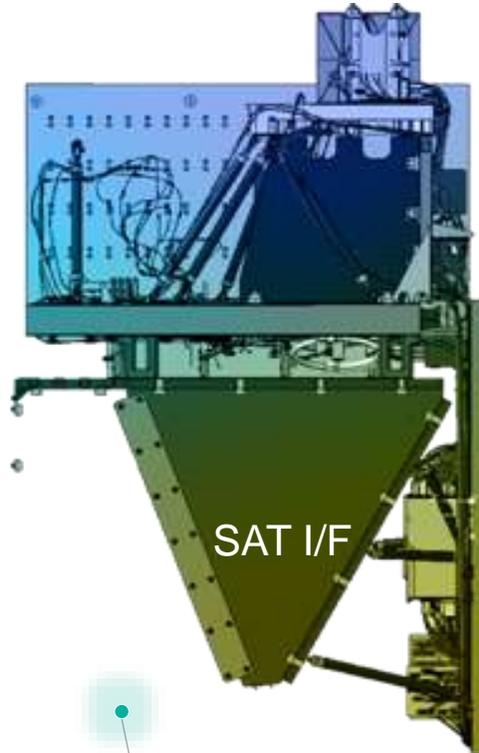
IASI-NG OVERALL DESIGN

- Afocal and Imaging Telescopes
- Mertz Interferometer
- **Focal Plane** : 4 detectors for 4 spectral bands /16 sounding pixels per detector
- **5 metrology lasers:**
 - 1 central metrology to give the Optical Path Difference constant triggering
 - 4 lateral lasers to monitor in real time the pupil effects (tilt, focus and astigmatism) for correction through on ground processing

SPECIFICATIONS			
GEOMETRY	SOUNDER PIXEL SIZE	~12 km	} SAME AS IASI
	SPATIAL SAMPLING	~25 km	
	GEOLOCATION ERROR	0.5 km	
SPECTRAL	BAND	645 cm ⁻¹ to 2760 cm ⁻¹	} 2 TIMES BETTER THAN IASI
	RESOLUTION	0,25 cm ⁻¹	
	SAMPLING	0,12 cm ⁻¹	
	CALIBRATION ERROR	da/σ= 5.10 ⁻⁷	
RADIOMETRY	CALIBRATION ERROR	0,25K @ 280 K	}
	NEDT	NedT -0.1 K to 0.4 K within spectrum	

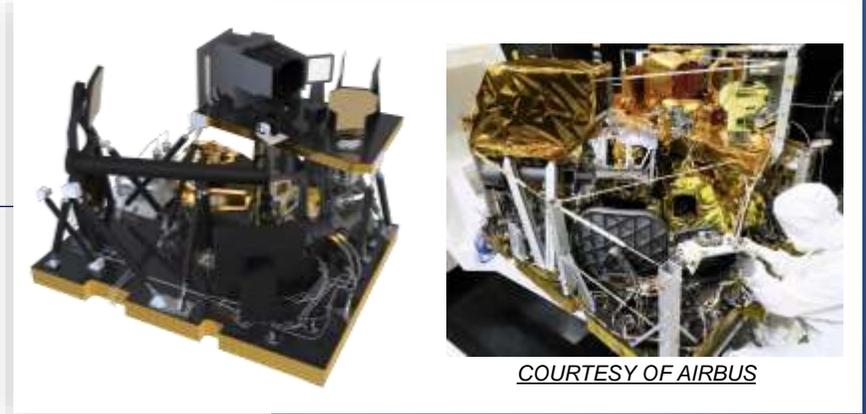
MAIN CHARACTERISTICS	
SWATH	~ 2000 KM
FOR	+/- 3°
PUPIL DIAMETER	~ 90 MM
ATA MAGNIFICATION	2.3
MAXIMUM OPTICAL PATH DIFFERENCE	4,2 CM
ACQUISITION DURATION	~730 MS
SCAN LINE DURATION	15.6 s : 14 EARTH VIEWS + 1BB + 1CS
CO-REGISTRATION	INTEGRATED IMAGER
SPECTRAL CALIBRATION	FABRY PEROT SOURCE

IASI-NG INSTRUMENT BUDGET



I-OH
(OPTICAL
MODULE)

I-EM
(ELECTRONICS
MODULE)



COURTESY OF AIRBUS

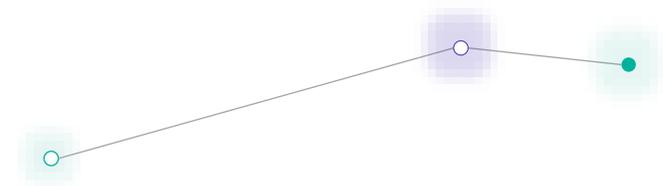


COURTESY OF AIRBUS

MASS	430 KG
POWER	~520 W
I/F DATA RATE	~6 MB/S

IASI-NG CURRENT CONTEXT

- Proto-Flight Model integrated and tested in 2022. Verification incomplete due to issues during TVAC, and an anomaly in B4 band was found
- PFM delivered and mounted on Satellite to give priority to qualification at Satellite level (mainly mechanical and thermal); and then dismantled for open work activities
- In the meantime, FM2 integration and tests were done. Confirmation that no B4 anomaly on FM2 and **formal decision to deliver FM2 as the first IASI-NG instrument for flight**
- Handover to Satellite held on 10th September 2024 and mechanical mounting performed on 12th September



IASI-NG

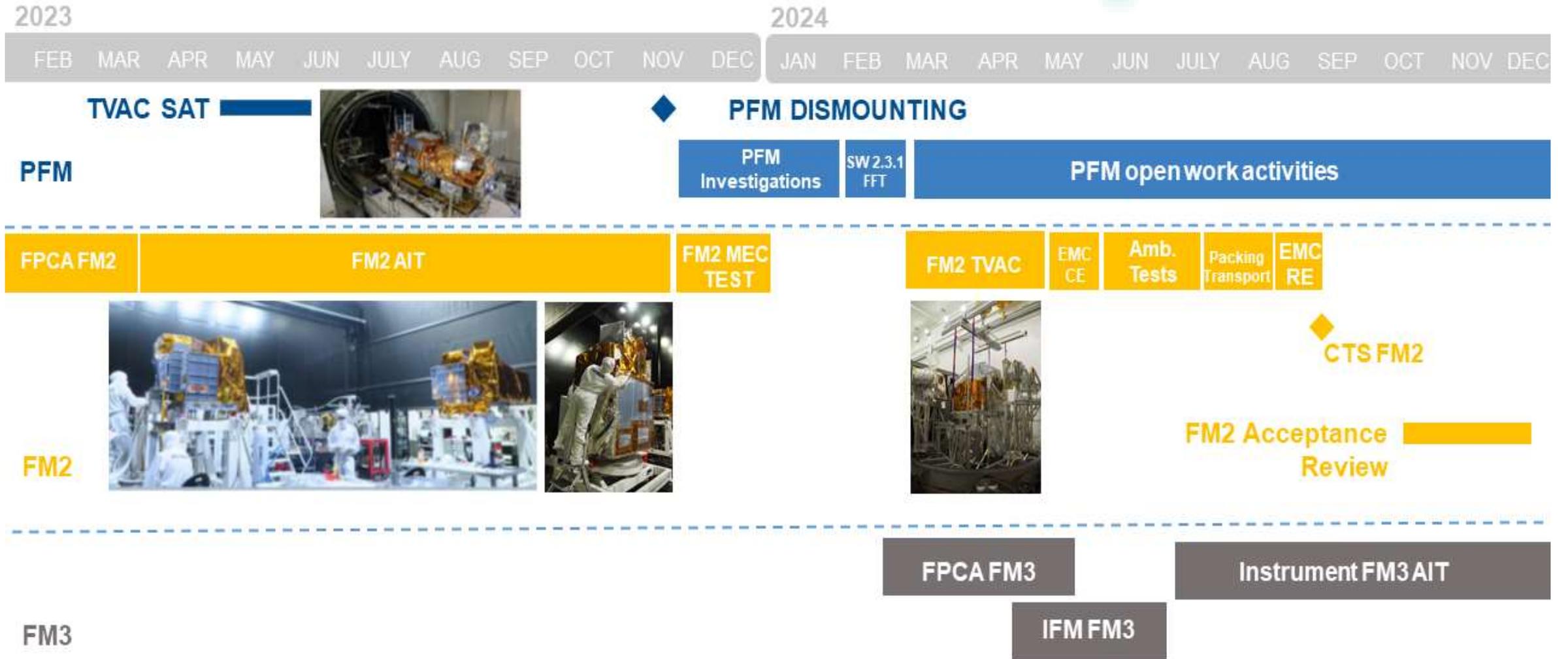


METOP SG-A1 ON SHAKER
(courtesy of AIRBUS)



METOP SG-A1 IN FRONT OF TV CHAMBER
(courtesy of AIRBUS)

IASI-NG CURRENT CONTEXT



IASI-NG FM2 PERFORMANCE VERIFICATION

- Performance verification is done in two steps
- TVAC for assessment of spectral and radiometric fine performances
 - 2 external warm black bodies and 1 cold black body for radiometric performances
 - 4 lasers for Instrument Spectral Response (+spectral calibration stability)
 - Gas Cell (+cold black body) for spectral calibration
 - Cold black body in front of Fabry Perot view
- Ambient test for geometric performances (LOS, IPSF, crosstalk)
 - Warm black body and collimator OGSE



FM2 INSTRUMENT IN TVAC CHAMBER
(COURTESY OF AIRBUS)



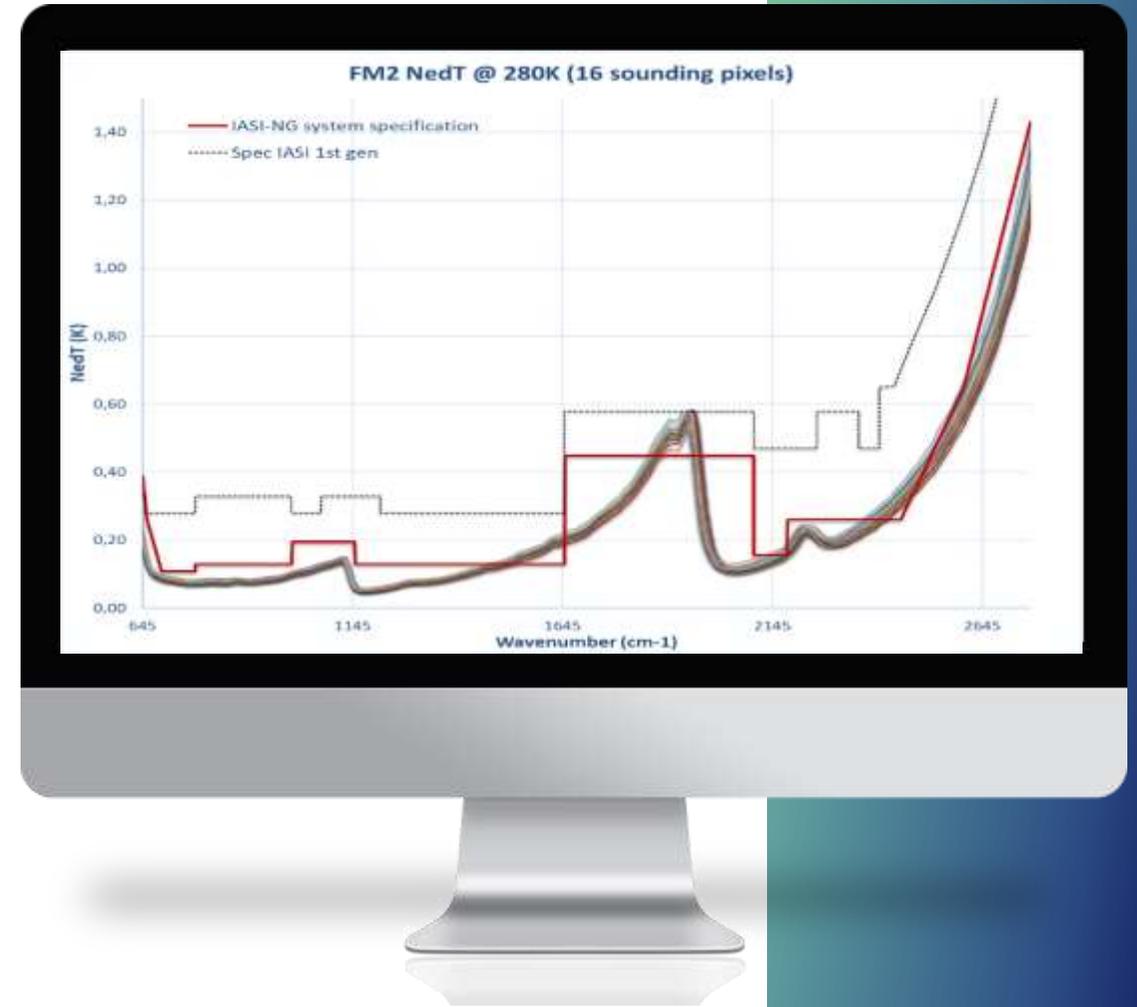
THERMAL TENT INSTALLATION
(COURTESY OF AIRBUS)



AMBIENT TEST SET-UP
(COURTESY OF AIRBUS)

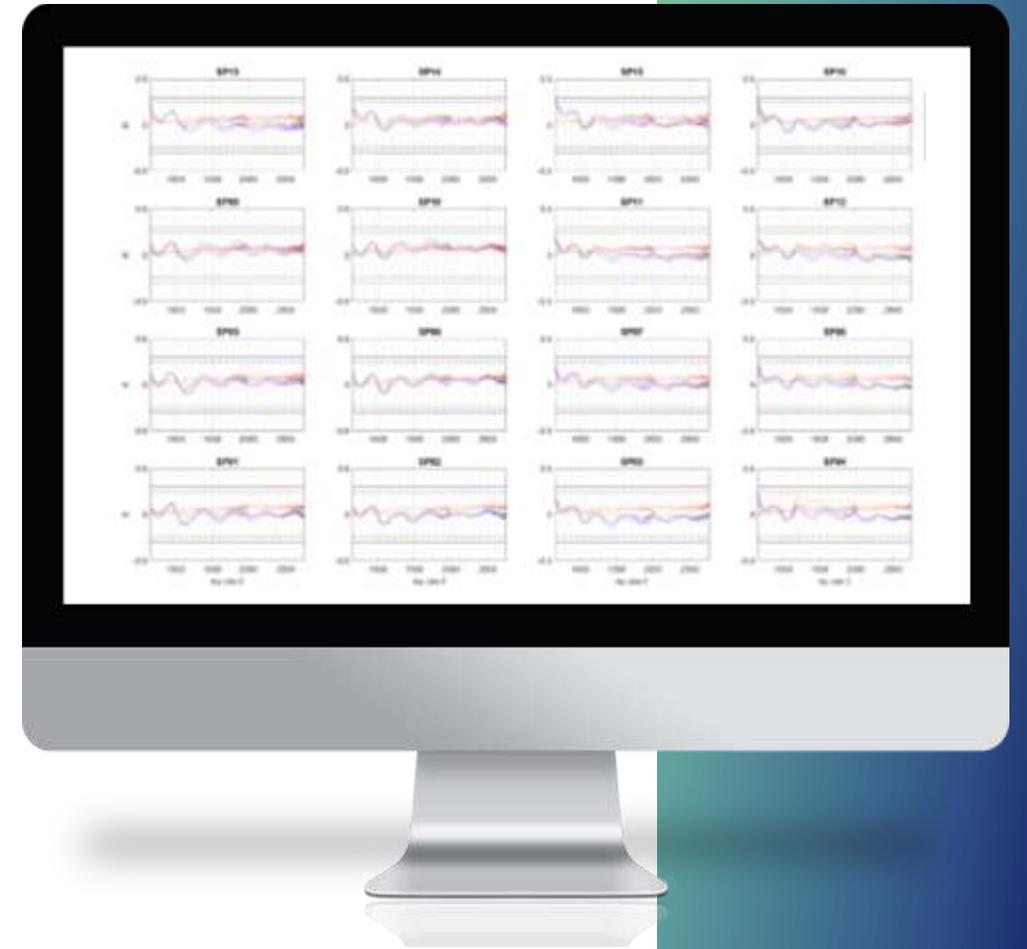
RADIOMETRIC PERFORMANCES

- NedT higher than CDR prediction due to higher background flux (in phase opposition compared to usefull signal)
- ✔ Nevertheless, NedT @ 280K compliant with system specification over the whole spectrum excepted in limited spectral area
- ✔ Confirmation of strong improvement with regard to IASI 1st generation specification (on top of the 2 times better spectral resolution)



RADIOMETRIC PERFORMANCES

- ✔ Absolute calibration error in line with specification ($\pm 0.25\text{K}$) for all tested cases (from 210K to 310K)
- Residual errors gives non compliances on inter-channel and inter-pixel calibration error due to two major phenomena:
 - vignetting effect on extreme views due to scan sizing issue with regard to pupil real size
 - difficulty to correct detector NL: due to high background level in phase opposition, internal BB view does not cover the whole useful dynamic
- These NC have been presented at ISSWG and are acceptable



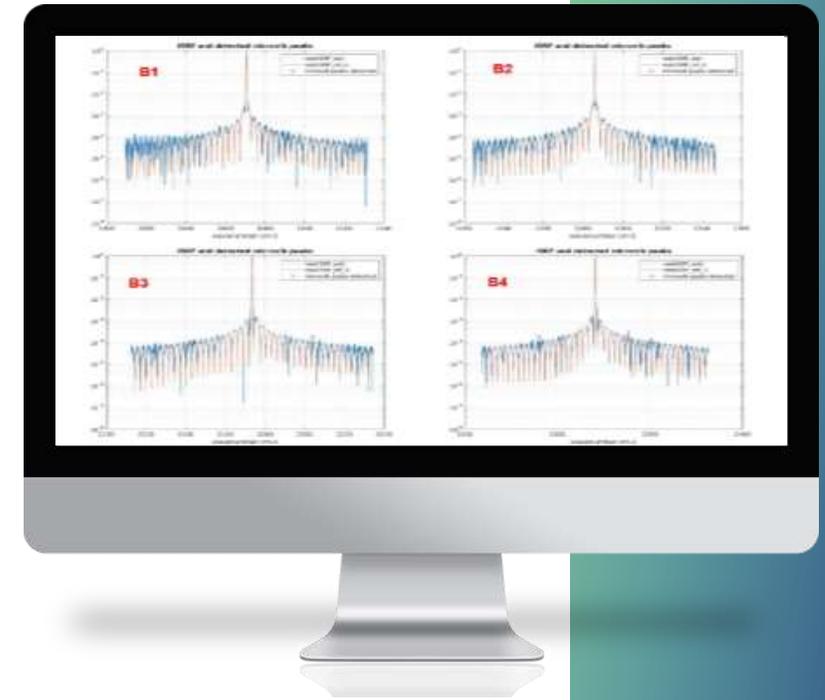
SPECTRAL PERFORMANCES

✓ ISRF with very good performances:

- Very low μ vibration peaks
- No significant parasitic interferogram peak
- Shape error index globally in line with CDR WC budget (this contributor is mainly a fixed one and induces a radiometric bias on atmospheric scenes)

• Improvement still possible by optimizing the model in the ground processing:

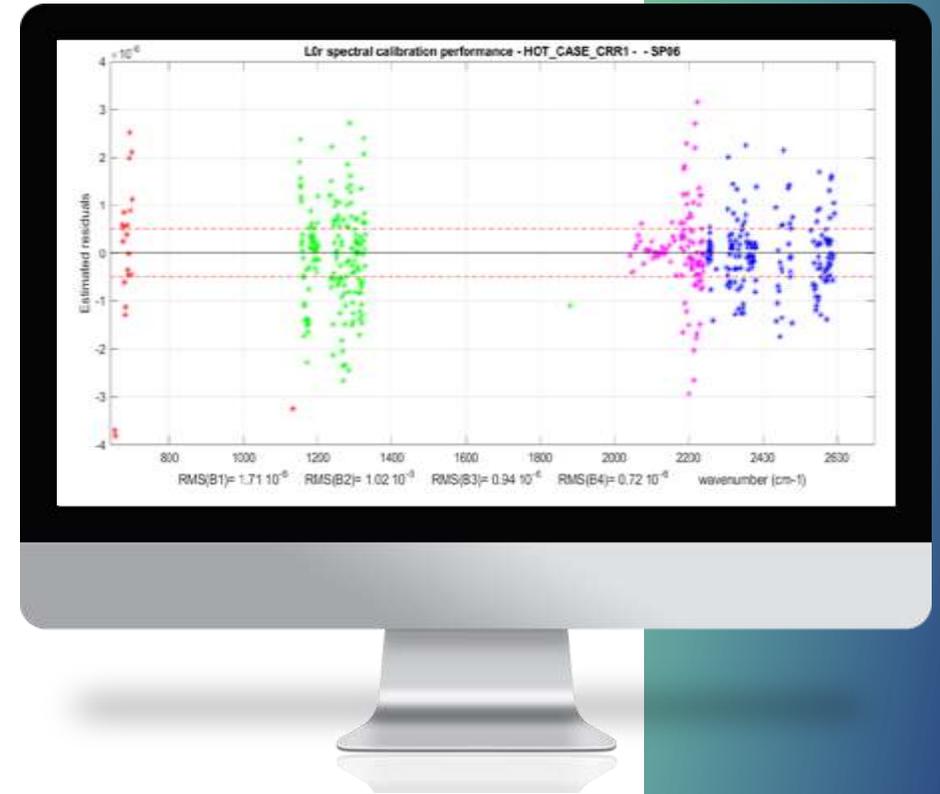
- Beam Splitter mechanism sine mode used during TVAC allows an optimization of the model parameters by checking the correction efficiency of induced tilt by the sinus vibration
- Could improve the shape error index



Performance	CDR WC budget	Measured Performance
Shape Error Index	B1: 1.7% B2: 1.8% B3: 2.5% B4: 3%	B1: 2.2% B2: 2.0% B3: 2.3% B4: 2.7%
Un-modelled parameters	0.39%	<0.10%
OPD μ vibrations	0.8% (0.5% RSS)	<0.33% (<0.2% RSS)
Contrast μ vibrations	0.5% (0.44% RSS)	<0.40% (<0.2% RSS)

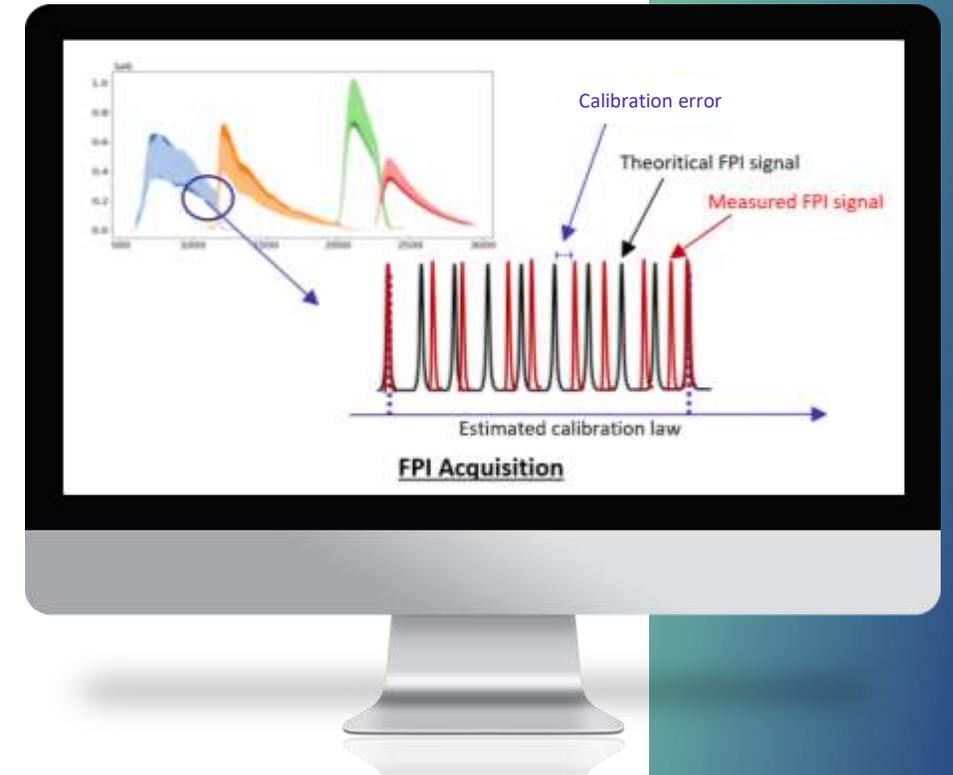
SPECTRAL PERFORMANCES: CALIBRATION

- Spectral calibration based on 6 parameters KBr index law applied over the full spectrum
 - Gas Cell used to fit the law on reference lines, and centroid shift (spectral calibration error) is estimated with verification lines :
 - Mean residual error is only few 10^{-7}
 - Rms residual errors is around 1 to $3 \cdot 10^{-6}$ in B1 in WC, due to Gas Cell line knowledge and lack of lines in B1
- ✓ This very good results with “all bands” approach confirm the validity of the spectral law modelling
- ✓ It shows the compliance to the “a priori” knowledge of the centroid shift (10^{-4})



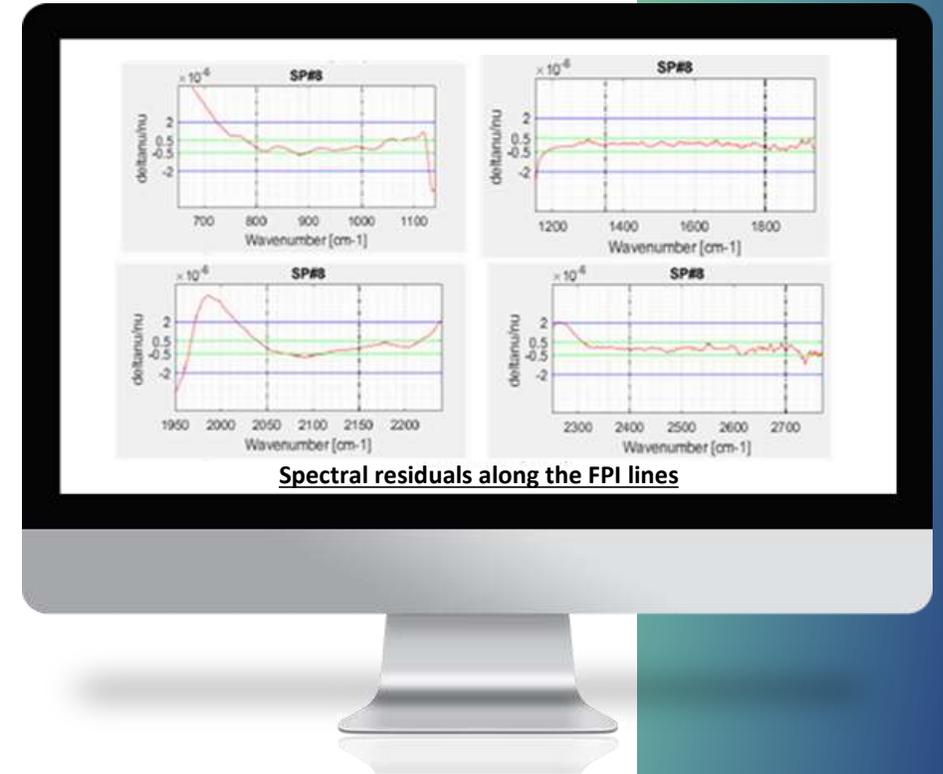
SPECTRAL PERFORMANCES: CALIBRATION

- The “a posteriori” knowledge specification supposing using perfectly known lines is $5 \cdot 10^{-7}$
- Verification by using gas cell lines is limited
- On board Fabry Perot Interferometer is used on this purpose, as the Free Spectral Range is almost constant



SPECTRAL PERFORMANCES: CALIBRATION

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- On board Fabry Perot Interferometer is used on this purpose, as the Free Spectral Range is almost constant
- ✔ **Residuals between two reference FPI peaks is very low and $<5.10^{-7}$**
- Residuals higher in inter-bands but suspected to be due to IPSF impact on FPI FSR (under analysis)
- It shows that residuals not directly linked to KBr index law are compliant with specification, and then that “a posteriori” specification is well achievable





**THANK YOU FOR
YOUR ATTENTION**

